

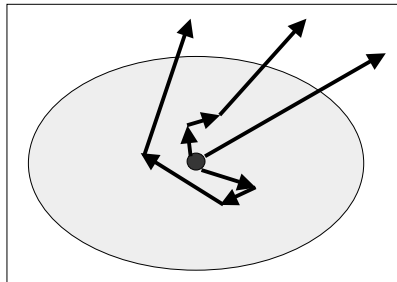
12. Circumstellar Matter

Monte Carlo Radiation Transfer I

- Monte Carlo “Photons” and interactions
- Sampling from probability distributions
- Optical depths, isotropic emission, scattering

Monte Carlo Basics

- Emit energy packet, hereafter a “photon”
- Photon travels some distance
- Something happens...



- Scattering, absorption, re-emission

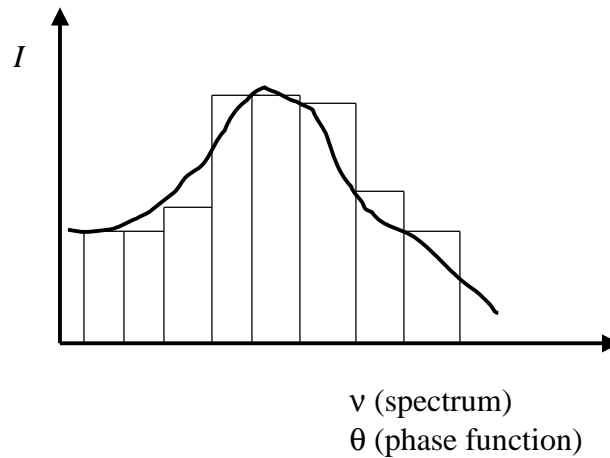
Photon Packets

If the total input luminosity is L , then each photon packet carries energy $E_i = L \Delta t / N$, where N is the number of Monte Carlo photons. A Monte Carlo photon represents N_γ real photons, where $N_\gamma = E_i / h\nu_i$. So, a Monte Carlo photon packet moving along direction specified by θ will contribute to the specific intensity:

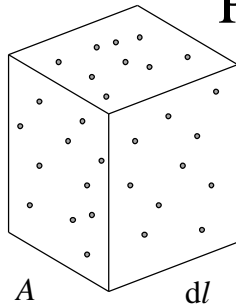
$$I_\nu = \frac{dE_\nu}{\cos \theta dA dt d\nu d\Omega}$$

$$\Delta I_\nu = \frac{E_i}{\cos \theta \Delta A \Delta t \Delta \nu \Delta \Omega} \longrightarrow \text{Energy packet}$$

Note, I_ν is a ***distribution function***. We will be working with ***discrete*** energies. Binning the photon packets into directions, frequencies, etc, enables us to simulate a distribution function. e.g., spectrum: bin in frequency; scattering phase function: bin in angle



Photon Interactions



Volume = $A \, dl$

Number density n

Cross section σ

Energy removed from beam per $t / v / d\Omega = I_v \sigma$

Number of photons absorbed/scattered from beam per sec
 $= I_v \sigma n A \, dl$

Number of photons absorbed/scattered from beam per sec per area
 $= I_v \sigma n \, dl$

Intensity differential over dl is $dI_v = -I_v n \sigma \, dl$. Therefore

$$I_v(l) = I_v(0) \exp(-n \sigma l)$$

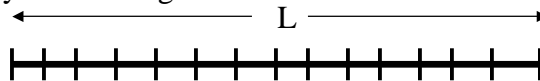
Fraction scattered or absorbed / length = $n \sigma$

$n \sigma$ = volume absorption coefficient = $\rho \kappa$

Mean free path = $1 / n \sigma$ = average distance between interactions

Probability of interaction over dl is $n \sigma \, dl$

Probability of traveling dl without interaction is $1 - n \sigma \, dl$



N segments of length L / N

Probability of traveling L before interacting is

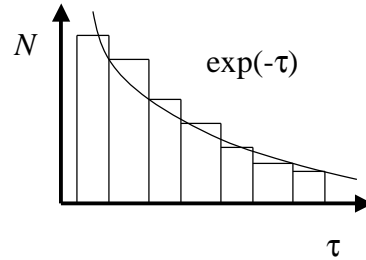
$$\begin{aligned} P(L) &= (1 - n \sigma l / N) (1 - n \sigma l / N) (1 - n \sigma l / N) \dots \\ &= (1 - n \sigma l / N)^N = \exp(-n \sigma L) \end{aligned}$$

$$P(L) = \exp(-\tau)$$

τ = number of mean free paths over distance L .

Probability Distribution Function

The probability distribution function (PDF) for photons to travel optical depth τ before an interaction is $\exp(-\tau)$. If we pick τ uniformly over the range 0 to infinity we will not reproduce $\exp(-\tau)$. We want to pick lots of small τ s and fewer large τ s. Same with a scattering phase function: want to get the correct number of photons scattered into different directions, forward and back scattering, etc.



Cumulative Distribution Function

$$\text{CDF} = \text{Area under PDF} = \int P(x) dx$$

Want to randomly choose τ , θ , λ , ... so that PDF is reproduced

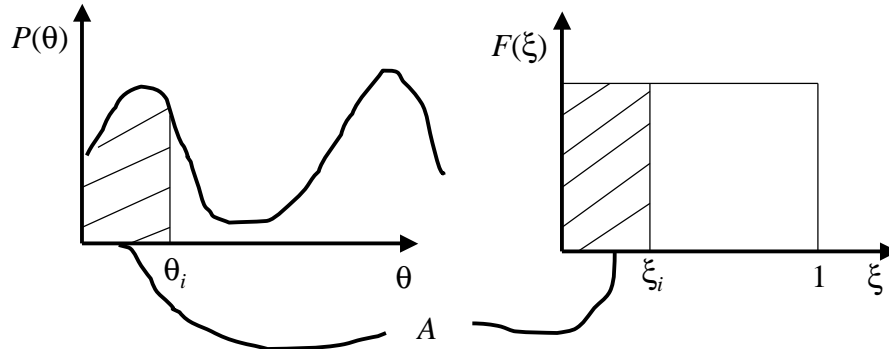
ξ is a random number
uniformly chosen in
range $[0,1]$

$$\xi = \int_0^x P(x) dx \Rightarrow X$$

$$\int_{-\infty}^{\infty} P(x) dx = 1$$

The above equation is the *fundamental principle* behind Monte Carlo techniques and is used to sample randomly from PDFs.

e.g., $P(\theta) = \cos \theta$ and we want to map ξ to θ . Choose random θ s to “fill in” $P(\theta)$



$$\xi_i = \int_0^{\theta_i} P(\theta) d\theta = \sin \theta_i \Rightarrow \theta_i = \sin^{-1} \xi_i$$

If we sample many random θ_i in this way and “bin” them, we will reproduce the curve $P(\theta) = \cos \theta$.

Choosing a Random Optical Depth

$P(\tau) = \exp(-\tau)$, i.e., photon travels τ before interaction

$$\xi = \int_0^{\tau} e^{-\tau} d\tau = 1 - e^{-\tau} \Rightarrow \tau = -\log(1 - \xi)$$

Since ξ is in range $[0,1]$, then $(1-\xi)$ is also in range $[0,1]$, so we may write:

$$\tau = -\log \xi$$

We find the physical distance, L , that the photon has traveled from:

$$\tau = \int_0^L n \sigma ds$$

Random Isotropic Direction

Solid angle is $d\Omega = \sin \theta \, d\theta \, d\phi$, we want to choose (θ, ϕ) so they fill in PDFs for θ and ϕ . $P(\theta)$ normalized over $[0, \pi]$, $P(\phi)$ normalized over $[0, 2\pi]$:

$$P(\theta) = \frac{1}{2} \sin \theta \quad P(\phi) = 1 / 2\pi$$

Using fundamental principle from above:

$$\xi = \int_0^\theta P(\theta) d\theta = \frac{1}{2} \int_0^\theta \sin \theta d\theta = \frac{1}{2} (\cos \theta - 1)$$

$$\xi = \int_0^\phi P(\phi) d\phi = \frac{1}{2\pi} \int_0^\phi d\phi = \frac{\phi - 1}{2\pi}$$

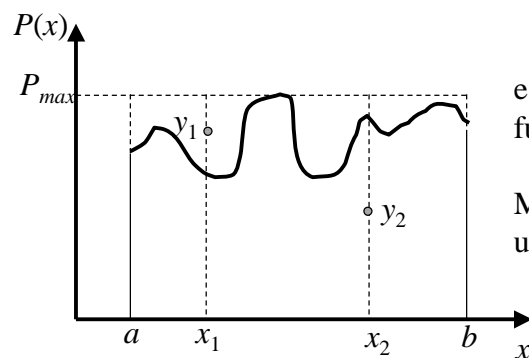
$$\theta = \cos^{-1}(2\xi - 1)$$

$$\phi = 2\pi \xi$$

Use this formula for emitting photons isotropically from a point source, or for choosing a scattering direction for isotropic scattering.

Rejection Method

The rejection method is used when we cannot invert the PDF as in the above examples to obtain analytic formulae for random θ , λ , etc.



e.g., $P(x)$ can be complicated function or tabulated

Multiply two random numbers: uniform probability / area

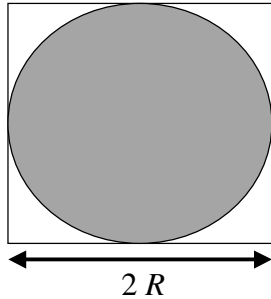
Choose x_1 in range $[a, b]$: $x_1 = a + \xi(b - a)$, calculate $P(x_1)$

Choose y_1 in range $[0, P_{max}]$: $y_1 = \xi P_{max}$

If $y_1 > P(x_1)$, reject x_1 . Choose new x_2, y_2 until $y_2 < P(x_2)$: accept x_2

Efficiency = Area under $P(x)$

Calculate π by the Rejection Method



FORTRAN 77:

Choose N random positions (x_i, y_i) :
 Choose x_i in range $[-R, R]$: $x_i = (2\xi - 1) R$
 Choose y_i in range $[-R, R]$: $y_i = (2\xi - 1) R$
 Reject (x_i, y_i) if $x_i^2 + y_i^2 > R^2$
 Number accepted / $N = \pi R^2 / 4R^2$
 $N_A / N = \pi / 4$
 Increase accuracy (signal to noise) by increasing N .

```
do i = 1, N
  x = 2.*ran - 1.
  y = 2.*ran - 1.
  if ( (x*x + y*y) .lt. 1. ) NA = NA + 1
end do
pi = 4.*NA / N
```

Albedo

When photon gets to interaction location at the randomly chosen optical depth, τ , we must decide whether the photon is scattered or absorbed. We use the *albedo* or *scattering probability*. It is the ratio of scattering to total opacity:

$$a = \frac{\sigma_s}{\sigma_s + \sigma_A}$$

To decide if a photon is scattered we choose a random number in the range $[0, 1]$ and scatter if $\xi < a$, otherwise the photon is absorbed.

We now have the tools required to write a Monte Carlo radiation transfer program for isotropic scattering in a constant density slab...